#### THE STANDARD MODEL AND EXPERIMENT

John Iliopoulos, ENS, Paris

Torino

February 8, 2019

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

#### The STANDARD MODEL in Particle Physics

#### A Quantum Field Theory describing in a unified framework

all

experimentally known interactions among elementary particles.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

#### What is an ELEMENTARY PARTICLE?

or, What is the World made of ?



#### What is an ELEMENTARY PARTICLE?

or, What is the World made of ?

#### I. The constituents of matter

Over the last century we have uncovered many layers of this cosmic onion:

```
atoms \rightarrow

nuclei + electrons \rightarrow

protons + neutrons + electrons \rightarrow

quarks + electrons \rightarrow ??
```

There is no reason to believe that there exists such a thing as "an innermost layer" and, even less, that we have already reached it.

### What is an ELEMENTARY PARTICLE?

#### or, What is the World made of ?

II. The Quanta of Radiation

Like the photon, they transmit the interactions. We know that all interactions are mediated by the exchange of such quanta.

The range of every interaction depends on the mass of the corresponding quantum

$$V(r) \sim \frac{\mathrm{e}^{-mr}}{r}$$

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

TABLE OF ELEMENTARY PARTICLES			
QUANTA OF RADIATION			
Strong Interactions		Eight gluons	
Electromagnetic Interactions		Photon $(\gamma)$	
Weak Interactions		Bosons $W^+$ , $W^-$ , $Z^0$	
Gravitational Interactions		Graviton (?)	
MATTER PARTICLES			
	Leptons	Quarks	
1st Family	$ u_{e}$ , $e^{-}$	$u_{a}$ , $d_{a}$ , $a=1,2,3$	
2nd Family	$ u_{\mu}$ , $\mu^-$	$c_a$ , $s_a$ , $a=1,2,3$	
3rd Family	$ u_{ au}$ , $ au^-$	$t_{a}$ , $b_{a}$ , $a=1,2,3$	
BROUT-ENGLERT-HIGGS BOSON			

This Table shows our present ideas on the structure of matter. Quarks and gluons do not exist as free particles and the graviton has not yet been observed

The strong interactions.
 Responsible for nuclear structure.

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

- The strong interactions.
   Responsible for nuclear structure.
- The electromagnetic interactions.
   Responsible for atomic and molecular structure.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- The strong interactions.
   Responsible for nuclear structure.
- The electromagnetic interactions.
   Responsible for atomic and molecular structure.
- The weak interactions.

Responsible for nuclear  $\beta\text{-decay}$  as well as the decays of other unstable particles.

(ロ) (型) (E) (E) (E) (O)

- The strong interactions.
   Responsible for nuclear structure.
- The electromagnetic interactions.
   Responsible for atomic and molecular structure.
- The weak interactions.

Responsible for nuclear  $\beta$ -decay as well as the decays of other unstable particles.

The gravitational interactions.

Manifest in everyday life, they are responsible for the large scale structure of the Universe. At the microscopic level, their effects are too small to be observable. The Standard Model describes accurately the strong, the electromagnetic and the weak interactions.

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

It is based on two fundamental principles:

The Standard Model describes accurately the strong, the electromagnetic and the weak interactions.

It is based on two fundamental principles:

• A Dynamical Theory  $\equiv$  A Quantum Field Theory

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

The Standard Model describes accurately the strong, the electromagnetic and the weak interactions.

It is based on two fundamental principles:

• A Dynamical Theory  $\equiv$  A Quantum Field Theory

 A property of Symmetry (a "gauge" symmetry) which brings Geometry into Physics.

ション ふゆ アメリア メリア しょうくしゃ

## I. The Dynamics

The two classical forces -Electromagnetism -Gravitation

are both described by the same classical potential:

 $V(r) \sim 1/r$ 

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

which is singular for  $r \rightarrow 0$ .

A classical atom is unstable!

 Non-Relativistic Quantum Mechanics solves this problem

$$\Delta(x) \ \Delta(p) \geq \hbar$$

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

The energy levels in an electromagnetic *or* a gravitational potential are quantised.

 Non-Relativistic Quantum Mechanics solves this problem

$$\Delta(x) \ \Delta(p) \geq \hbar$$

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

The energy levels in an electromagnetic *or* a gravitational potential are quantised.

but the relativistic corrections bring it back!

 Non-Relativistic Quantum Mechanics solves this problem

$$\Delta(x) \ \Delta(p) \geq \hbar$$

The energy levels in an electromagnetic *or* a gravitational potential are quantised.

- but the relativistic corrections bring it back!
- A remark: The uncertainty relations solve the problem of the 1/r potential.
   Not for every potential

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

Birth of Quantum Electrodynamics

The first consistent Quantum Field Theory, free of singularities at all distances!

Birth of Quantum Electrodynamics

The first consistent Quantum Field Theory, free of singularities at all distances!

Quantum Field Theories satisfy also uncertainty relations

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

Birth of Quantum Electrodynamics

The first consistent Quantum Field Theory, free of singularities at all distances!

- Quantum Field Theories satisfy also uncertainty relations
- Quantum mechanics uncertainty relations do NOT solve the problem for all potentials

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

Birth of Quantum Electrodynamics

The first consistent Quantum Field Theory, free of singularities at all distances!

- Quantum Field Theories satisfy also uncertainty relations
- Quantum mechanics uncertainty relations do NOT solve the problem for all potentials
- Quantum field theory uncertainty relations do NOT solve the problem for all quantum field theories. (Renormalisable QFTs)

Birth of Quantum Electrodynamics

The first consistent Quantum Field Theory, free of singularities at all distances!

- Quantum Field Theories satisfy also uncertainty relations
- Quantum mechanics uncertainty relations do NOT solve the problem for all potentials
- Quantum field theory uncertainty relations do NOT solve the problem for all quantum field theories. (Renormalisable QFTs)
- Quantum Electrodynamics seemed to be the only interesting case

Classical Mechanics q(t), p(t)	$\Rightarrow$ [q(t), p(t)] = $i\hbar$ $\Rightarrow$	Quantum Mechanics
$\psi \\ q_i(t), p_i(t) \\ i = 1,, N$	$\Rightarrow [q_i(t), p_i(t)] = i\hbar\delta_{ij}$	$\Downarrow$
$N  ightarrow \infty$		
$\Downarrow$		$\Downarrow$
Classical Field $\Rightarrow$ [a Theory $q(\vec{x}, t), p(\vec{x}, t)$	$p(ec{x},t),p(ec{y},t)]=i\hbar\delta^3(ec{x}-ec{y})$ =	Quantum → Field Theory

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ④�?

# Classical Field Theory = Classical Mechanics

With an infinite number of degrees of freedom



# Classical Field Theory = Classical Mechanics

With an infinite number of degrees of freedom

## Quantum Field Theory

# Quantum Mechanics

With an infinite number of degrees of freedom

◆□▶ ◆□▶ ★□▶ ★□▶ □ のQ@

# Classical Field Theory = Classical Mechanics

With an infinite number of degrees of freedom

#### Quantum Field Theory

# Quantum Mechanics

With an infinite number of degrees of freedom

It is always the case with a relativistic theory

(ロ) (型) (E) (E) (E) (O)

▲□▶ ▲圖▶ ▲国▶ ▲国▶ 三国 - のへで

▶ 1925 : Born-Heisenberg-Jordan

The free electromagnetic field as an infinite set of harmonic oscillators

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

▶ 1925 : Born-Heisenberg-Jordan

The free electromagnetic field as an infinite set of harmonic oscillators

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

▶ 1927 : Dirac

The spontaneous emission probability

▶ 1925 : Born-Heisenberg-Jordan

The free electromagnetic field as an infinite set of harmonic oscillators

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

▶ 1927 : Dirac

The spontaneous emission probability

▶ 1930 : Ambartsumian-Ivanenko

Canonical quantisation of massive fields

▶ 1925 : Born-Heisenberg-Jordan

The free electromagnetic field as an infinite set of harmonic oscillators

▶ 1927 : Dirac

The spontaneous emission probability

► 1930 : Ambartsumian-Ivanenko

Canonical quantisation of massive fields

▶ 1933 : Perrin

In nuclear  $\beta\text{-decay},$  the emitted electrons and neutrinos are created the moment of emission

(ロ) (型) (E) (E) (E) (O)

▶ 1925 : Born-Heisenberg-Jordan

The free electromagnetic field as an infinite set of harmonic oscillators

▶ 1927 : Dirac

The spontaneous emission probability

▶ 1930 : Ambartsumian-Ivanenko

Canonical quantisation of massive fields

▶ 1933 : Perrin

In nuclear  $\beta\text{-decay},$  the emitted electrons and neutrinos are created the moment of emission

▶ 1933 : Fermi

Fermions quantised. Quantum Field Theory becomes the language of microscopic physics

#### Renormalisable theories

In our four dimensional space there exist FIVE renormalisable quantum field theories:

- $\phi^3(x)$
- $\phi^4(x)$
- The Yukawa interaction:  $\bar{\psi}(x)\psi(x)\phi(x)$
- QED:  $\bar{\psi}(x)\gamma_{\mu}\psi(x)A^{\mu}(x)$ (and scalar QED)
- The Yang-Mills interaction:  $Tr(F^{\mu\nu}(x)F_{\mu\nu}(x))$

ション ふゆ く 山 マ チャット しょうくしゃ

#### Remarks

▲□▶ ▲圖▶ ▲国▶ ▲国▶ 目 のへで

#### Remarks

 Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson

#### Remarks

- Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson
- Renormalisable theories, like 1/r potentials, are not very sensitive to, necessarily unknown, physics at arbitrarily short distances

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson
- Renormalisable theories, like 1/r potentials, are not very sensitive to, necessarily unknown, physics at arbitrarily short distances

ション ふゆ アメリア メリア しょうくの

The gravitational interaction is not one of them

- Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson
- Renormalisable theories, like 1/r potentials, are not very sensitive to, necessarily unknown, physics at arbitrarily short distances
- The gravitational interaction is not one of them
- Nature uses ALL five renormalisable theories, and ONLY them, as fundamental theories

ション ふゆ く 山 マ チャット しょうくしゃ

- Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson
- Renormalisable theories, like 1/r potentials, are not very sensitive to, necessarily unknown, physics at arbitrarily short distances
- The gravitational interaction is not one of them
- Nature uses ALL five renormalisable theories, and ONLY them, as fundamental theories

ション ふゆ く 山 マ チャット しょうくしゃ

We only have approximate solutions

- Our modern views on Quantum Field Theory have been profoundly reshaped by K. Wilson
- Renormalisable theories, like 1/r potentials, are not very sensitive to, necessarily unknown, physics at arbitrarily short distances
- The gravitational interaction is not one of them
- Nature uses ALL five renormalisable theories, and ONLY them, as fundamental theories
- We only have approximate solutions
- The effective strength of the interaction depends in a calculable way on the energy, or distance, scale. (Renormalisation group)

It covers the entire domain of fundamental physics

▲□▶ ▲圖▶ ▲ 臣▶ ▲ 臣▶ ― 臣 … のへぐ

It covers the entire domain of fundamental physics

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

High Energy Physics

It covers the entire domain of fundamental physics

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

- High Energy Physics
- Nuclear structure

It covers the entire domain of fundamental physics

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- High Energy Physics
- Nuclear structure
- Condensed matter physics

It covers the entire domain of fundamental physics

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- High Energy Physics
- Nuclear structure
- Condensed matter physics
- Phase transitions and critical phenomena

It covers the entire domain of fundamental physics

ション ふゆ アメリア メリア しょうくの

- High Energy Physics
- Nuclear structure
- Condensed matter physics
- Phase transitions and critical phenomena
- Cosmology and astrophysics

II. Symmetries: A well known, but quite abstract concept

Symmetry = An assumption that a certain quantity has no physical meaning.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

Physically meaningful results cannot depend on it.

II. Symmetries: A well known, but quite abstract concept

Symmetry = An assumption that a certain quantity has no physical meaning.

Physically meaningful results cannot depend on it.

► The position of the origin of the coordinate system ⇒ We can move the coordinate system ⇒ Invariance under translations

II. Symmetries: A well known, but quite abstract concept

Symmetry = An assumption that a certain quantity has no physical meaning.

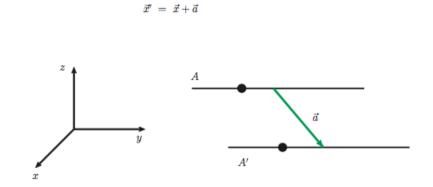
Physically meaningful results cannot depend on it.

► The position of the origin of the coordinate system ⇒ We can move the coordinate system ⇒ Invariance under translations

ション ふゆ く 山 マ チャット しょうくしゃ

► The direction of the axes in space ⇒ We can rotate the coordinate system ⇒ Invariance under rotations

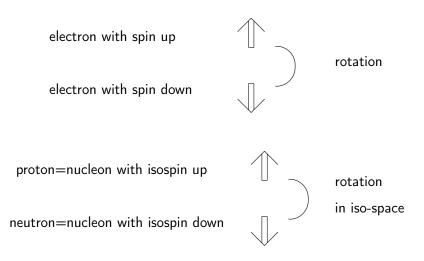
#### Invariance under translations



If A is the trajectory of a free particle in the (x,y,z) system, its image, A', is also a possible trajectory of a free particle.

#### The first abstraction: Internal Symmetries

Heisenberg 1932



・ロッ ・雪 ・ ・ ヨ ・ ・

э.

 Heisenberg's iso-space is three dimensional, isomorphic to our physical space.

- Heisenberg's iso-space is three dimensional, isomorphic to our physical space.
- With the discovery of new internal symmetries the idea was generalised to multi-dimensional internal spaces.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- Heisenberg's iso-space is three dimensional, isomorphic to our physical space.
- With the discovery of new internal symmetries the idea was generalised to multi-dimensional internal spaces.
- The space of Physics became an abstract mathematical concept with non-trivial geometrical and topological properties.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- Heisenberg's iso-space is three dimensional, isomorphic to our physical space.
- With the discovery of new internal symmetries the idea was generalised to multi-dimensional internal spaces.
- The space of Physics became an abstract mathematical concept with non-trivial geometrical and topological properties.

ション ふゆ く 山 マ チャット しょうくしゃ

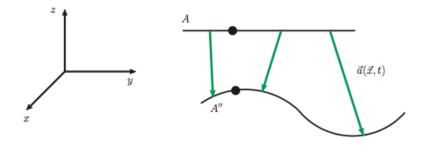
 Only a part of it, the three-dimensional Euclidean space, is directly accessible to our senses.

#### A further abstraction: Local Symmetries

Einstein 1918

Local space translations

 $\vec{x}'' = \vec{x} + \vec{a}(\vec{x}, t)$ 



▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

#### Local space translations

The question is purely geometrical without any obvious physical meaning, so we expect a mathematical answer with no interest for Physics.

・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・

#### Local space translations

The question is purely geometrical without any obvious physical meaning, so we expect a mathematical answer with no interest for Physics.

 Surprise: The Dynamics which is invariant under local translations is

#### GENERAL RELATIVITY

The resulting force is Gravity One of the four fundamental forces.

The gravitational forces are not the only ones which have a geometrical origin

The example of the quantum mechanical phase:

 $\Psi(x) 
ightarrow e^{i heta} \Psi(x)$  with heta 
ightarrow heta(x)

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

The gravitational forces are not the only ones which have a geometrical origin

The example of the quantum mechanical phase:

 $\Psi(x) \to e^{i\theta}\Psi(x) \quad \text{with} \quad \theta \to \theta(x)$  $\bullet \ \partial_{\mu}e^{i\theta(x)}\Psi(x) = e^{i\theta(x)}\partial_{\mu}\Psi(x) + ie^{i\theta(x)}\Psi(x)\partial_{\mu}\theta(x)$ 

The gravitational forces are not the only ones which have a geometrical origin

The example of the quantum mechanical phase:

 $\Psi(x) o e^{i\theta} \Psi(x)$  with heta o heta(x)

 $\bullet \ \partial_{\mu}e^{i\theta(x)}\Psi(x) = e^{i\theta(x)}\partial_{\mu}\Psi(x) + ie^{i\theta(x)}\Psi(x)\partial_{\mu}\theta(x)$ 

► Introduce  $A_{\mu}(x)$  such that  $A_{\mu}(x) \rightarrow A_{\mu}(x) - \frac{1}{e} \partial_{\mu} \theta(x)$ 

The gravitational forces are not the only ones which have a geometrical origin

The example of the quantum mechanical phase:

 $\Psi(x) \to e^{i\theta} \Psi(x)$  with  $\theta \to \theta(x)$ 

• Introduce  $A_{\mu}(x)$  such that

 $A_{\mu}(x) \rightarrow A_{\mu}(x) - \frac{1}{e}\partial_{\mu}\theta(x)$ 

Then

 $\partial_\mu o D_\mu = \partial_\mu - i e A_\mu(x)$  ;  $D_\mu e^{i heta(x)} \Psi(x) = e^{i heta(x)} D_\mu \Psi(x)$ 

ション ふゆ く 山 マ チャット しょうくしゃ

The gravitational forces are not the only ones which have a geometrical origin

- The example of the quantum mechanical phase:
  - $\Psi(x) o e^{i heta} \Psi(x) \qquad ext{with} \qquad heta o heta(x)$
- $\bullet \ \partial_{\mu}e^{i\theta(x)}\Psi(x) = e^{i\theta(x)}\partial_{\mu}\Psi(x) + ie^{i\theta(x)}\Psi(x)\partial_{\mu}\theta(x)$
- Introduce  $A_{\mu}(x)$  such that

 $A_{\mu}(x) \rightarrow A_{\mu}(x) - \frac{1}{e}\partial_{\mu}\theta(x)$ 

Then

 $\partial_\mu o D_\mu = \partial_\mu - i e A_\mu(x)$  ;  $D_\mu e^{i heta(x)} \Psi(x) = e^{i heta(x)} D_\mu \Psi(x)$ 

▶ Replacing ∂<sub>µ</sub> by D<sub>µ</sub> turns any equation which was invariant under the global phase transformation, invariant under the local (gauge) one. Fock 1926

► The introduction of the covariant derivative: The free Schrödinger, or Dirac, equation ⇒ The same equation in the presence of an external electromagnetic field

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

► The introduction of the covariant derivative:

The free Schrödinger, or Dirac, equation  $\Rightarrow$ 

The same equation in the presence of an external electromagnetic field

To obtain the fully interacting theory:

Add the energy of the new vector field:

 $\sim F_{\mu\nu}^2 = (\partial_\mu A_
u - \partial_
u A_\mu)^2$ 

The resulting interaction is:

#### QUANTUM ELECTRODYNAMICS

 The great progress of the last fifty years in our understanding of the fundamental forces was the realisation that They all obey the principle of Geometry

・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・

 The great progress of the last fifty years in our understanding of the fundamental forces was the realisation that They all obey the principle of Geometry

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

► They are all *gauge* forces

- The great progress of the last fifty years in our understanding of the fundamental forces was the realisation that They all obey the principle of Geometry
- ► They are all *gauge* forces

#### DYNAMICS = GEOMETRY

Platon : "Μηδείς αγεωμέτρητος εισίτω την στέγην."

"Let no one ignorant of geometry enter under this roof"

#### THE STANDARD MODEL

It describes, in a unified framework, the strong, the e.m. and the weak interactions.

#### THE STANDARD MODEL

It describes, in a unified framework, the strong, the e.m. and the weak interactions.

ション ふゆ アメリア メリア しょうくの

► It is based on a rather complicated gauge symmetry:  $U(1) \times SU(2) \times SU(3) \rightarrow U(1)_{em} \times SU(3)$ 

#### THE STANDARD MODEL

- It describes, in a unified framework, the strong, the e.m. and the weak interactions.
- ► It is based on a rather complicated gauge symmetry:  $U(1) \times SU(2) \times SU(3) \rightarrow U(1)_{em} \times SU(3)$
- It can be implemented in an internal space with as many as ten dimensions

ション ふゆ アメリア メリア しょうくの

#### THE STANDARD MODEL

- It describes, in a unified framework, the strong, the e.m. and the weak interactions.
- ► It is based on a rather complicated gauge symmetry:  $U(1) \times SU(2) \times SU(3) \rightarrow U(1)_{em} \times SU(3)$
- It can be implemented in an internal space with as many as ten dimensions

ション ふゆ アメリア メリア しょうくしゃ

But then we faced a new problem:
 THE PROBLEM OF MASS

No terms proportional to A<sub>µ</sub>A<sup>µ</sup>
 ⇒ the gauge fields describe massless particles.
 ⇒ Gauge symmetries imply long range correlations
 Useless for Physics??

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

No terms proportional to A<sub>µ</sub>A<sup>µ</sup>
 ⇒ the gauge fields describe massless particles.
 ⇒ Gauge symmetries imply long range correlations Useless for Physics??

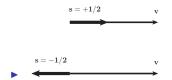
> The mass of the constituents of matter - chirality

ション ふゆ く 山 マ チャット しょうくしゃ

No terms proportional to A<sub>µ</sub>A<sup>µ</sup>
 ⇒ the gauge fields describe massless particles.
 ⇒ Gauge symmetries imply long range correlations Useless for Physics??

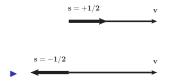
► The mass of the constituents of matter - chirality

◆□▶ ◆□▶ ◆□▶ ◆□▶ ● ● ●



No terms proportional to A<sub>µ</sub>A<sup>µ</sup>
 ⇒ the gauge fields describe massless particles.
 ⇒ Gauge symmetries imply long range correlations Useless for Physics??

► The mass of the constituents of matter - chirality

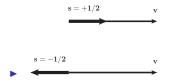


► Weak interactions involve only particles with one chirality ⇒ The constituents of matter must be massless

ション ふゆ く 山 マ チャット しょうくしゃ

No terms proportional to A<sub>µ</sub>A<sup>µ</sup>
 ⇒ the gauge fields describe massless particles.
 ⇒ Gauge symmetries imply long range correlations Useless for Physics??

► The mass of the constituents of matter - chirality



- ► Weak interactions involve only particles with one chirality ⇒ The constituents of matter must be massless
- Most of the particles in our Table are massive

#### This was THE PROBLEM OF MASS

# Spontaneous Symmetry Breaking (SSB)

 An infinite system may exhibit the phenomenon of phase transitions. It often implies a reduction in the symmetry of the ground state.

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

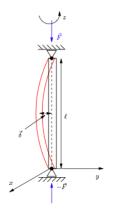
## Spontaneous Symmetry Breaking (SSB)

- An infinite system may exhibit the phenomenon of phase transitions. It often implies a reduction in the symmetry of the ground state.
- For a field theory, in many cases, we encounter at least two phases:

(i) The unbroken, or, the Wigner phase: A symmetry is manifest in the spectrum of the theory whose excitations form irreducible representations of the symmetry group. For a gauge theory the vector gauge bosons are massless and belong to the adjoint representation.

(ii) *The spontaneously broken phase*: Part of the symmetry is hidden from the spectrum. For a gauge theory, some of the gauge bosons become massive.

An example from Classical Mechanics



 $IE\frac{d^{4}X}{dz^{4}} + F\frac{d^{2}X}{dz^{2}} = 0 \quad ; \quad IE\frac{d^{4}Y}{dz^{4}} + F\frac{d^{2}Y}{dz^{2}} = 0$ X = X'' = Y = Y'' = 0 for z = 0 and z = I

An example from Classical Mechanics

• A symmetric solution always exists: X = Y = 0

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

An example from Classical Mechanics

• A symmetric solution always exists: X = Y = 0

► For 
$$F \ge F_{cr} = \frac{\pi^2 E I}{I^2}$$
 asymmetric solutions appear:  
 $X = C \sin kz$ ;  $kI = n\pi$ ;  $n = 1, ...$ ;  $k^2 = F/EI$ 

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

They correspond to lower energy.

An example from Classical Mechanics

• A symmetric solution always exists: X = Y = 0

• For 
$$F \ge F_{cr} = \frac{\pi^2 EI}{I^2}$$
 asymmetric solutions appear:  
 $X = C \sin kz$ ;  $kI = n\pi$ ;  $n = 1, ...$ ;  $k^2 = F/EI$   
They correspond to lower energy.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

What happened to the original symmetry?

An example from Classical Mechanics

- A symmetric solution always exists: X = Y = 0
- ► For  $F \ge F_{cr} = \frac{\pi^2 EI}{I^2}$  asymmetric solutions appear:  $X = C \sin kz$ ;  $kI = n\pi$ ; n = 1, ...;  $k^2 = F/EI$ They correspond to lower energy.

- What happened to the original symmetry?
- The ground state is degenerate.

An example from Classical Mechanics

- A symmetric solution always exists: X = Y = 0
- For  $F \ge F_{cr} = \frac{\pi^2 E l}{l^2}$  asymmetric solutions appear:

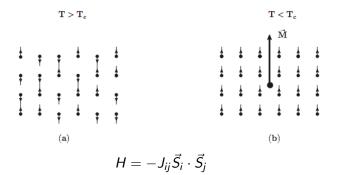
 $X = C \sin kz$ ;  $kl = n\pi$ ; n = 1, ...;  $k^2 = F/El$ 

They correspond to lower energy.

- What happened to the original symmetry?
- The ground state is degenerate.
- The state is characterised by the two-component vector  $\vec{\delta} = |\vec{\delta}|e^{i\theta}$ The modulus does have a physical meaning, the phase does not.

An example from Quantum Mechanics

The Heisenberg ferromagnet



◆□▶ ◆□▶ ◆□▶ ◆□▶ ● ● ●

Symmetry breaking  $O(3) \rightarrow O(2)$ 

A field theory example

• 
$$\mathcal{L}_1 = (\partial_\mu \phi)(\partial^\mu \phi^*) - M^2 \phi \phi^* - \lambda (\phi \phi^*)^2$$

Invariant under U(1) global transformations:  $\phi(x) \rightarrow e^{i\theta}\phi(x)$ 

• The Hamiltonian is given by:

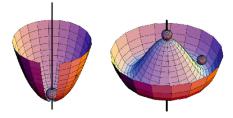
$$\begin{aligned} \mathcal{H}_1 &= (\partial_0 \phi)(\partial_0 \phi^*) + (\partial_i \phi)(\partial_i \phi^*) + V(\phi) \\ V(\phi) &= M^2 \phi \phi^* + \lambda (\phi \phi^*)^2 \end{aligned}$$

- The symmetric solution is  $\phi(x) = 0$ .
- The minimum energy configuration corresponds to:

 $\phi(x) = \text{constant} = \phi$  such that  $V(\phi)$  is minimum, solution of: V' = 0

ション ふゆ アメリア メリア しょうくしゃ

A field theory example



• The potential  $V(\phi)$  with  $\lambda > 0$  and  $M^2 \ge 0$  (left).

The only solution is the symmetric one  $\phi = 0$ .

• The potential  $V(\phi)$  with  $\lambda > 0$  and  $M^2 < 0$  (right).

 $\phi = 0$  is a local maximum. An entire circle of minima at the complex  $\phi$ -plane with radius  $v = (-M^2/2\lambda)^{1/2}$ . Any point on it corresponds to a spontaneous breaking of the U(1) symmetry.

A field theory example

• Conclusion:  $M^2 = 0$  is a critical point.

For  $M^2 > 0$  the symmetric solution is stable.

For  $M^2 < 0$  spontaneous symmetry breaking occurs.

◆□▶ ◆圖▶ ◆臣▶ ◆臣▶ 三臣 - のへぐ

A field theory example

• Conclusion:  $M^2 = 0$  is a critical point.

For  $M^2 > 0$  the symmetric solution is stable.

For  $M^2 < 0$  spontaneous symmetry breaking occurs.

• In order to reach the stable solution we translate the field  $\phi$ .  $\phi(x) = \frac{1}{\sqrt{2}} \left[ v + \psi(x) + i\chi(x) \right]$  $\mathcal{L}_1(\phi) \rightarrow \mathcal{L}_2(\psi, \chi) = \frac{1}{2} (\partial_\mu \psi)^2 + \frac{1}{2} (\partial_\mu \chi)^2 - \frac{1}{2} (2\lambda v^2) \psi^2$ 

$$\mathcal{L}_2(\psi, \chi) \equiv \frac{1}{2} (\partial_\mu \psi)^2 + \frac{1}{2} (\partial_\mu \chi)^2 - \frac{1}{2} (2\lambda V)^2 \psi^2$$
$$- \lambda V \psi (\psi^2 + \chi^2) - \frac{\lambda}{4} (\psi^2 + \chi^2)^2$$

A field theory example

• Conclusion:  $M^2 = 0$  is a critical point.

For  $M^2 > 0$  the symmetric solution is stable.

For  $M^2 < 0$  spontaneous symmetry breaking occurs.

• In order to reach the stable solution we translate the field  $\phi$ .  $\phi(x) = \frac{1}{\sqrt{2}} \left[ v + \psi(x) + i\chi(x) \right]$   $\mathcal{L}_1(\phi) \rightarrow \mathcal{L}_2(\psi, \chi) = \frac{1}{2} (\partial_\mu \psi)^2 + \frac{1}{2} (\partial_\mu \chi)^2 - \frac{1}{2} (2\lambda v^2) \psi^2$  $-\lambda v \psi (\psi^2 + \chi^2) - \frac{\lambda}{4} (\psi^2 + \chi^2)^2$ 

χ is massless (Goldstone mode).

A field theory example

•  $\mathcal{L}_2$  is still invariant.

$$\delta\psi = -\theta\chi$$
 ;  $\delta\chi = \theta\psi + \theta\mathbf{v}$ 

We still have a conserved current:

$$j_{\mu} \sim \psi \partial_{\mu} \chi - \chi \partial_{\mu} \psi + \mathbf{v} \partial_{\mu} \chi$$
  
 $\partial^{\mu} j_{\mu}(\mathbf{x}) = \mathbf{0}$ 

It is the minimum energy configuration which is not invariant.

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

A field theory example

•  $\mathcal{L}_2$  is still invariant.

$$\delta\psi = -\theta\chi$$
 ;  $\delta\chi = \theta\psi + \theta\mathbf{v}$ 

We still have a conserved current:

$$j_{\mu} \sim \psi \partial_{\mu} \chi - \chi \partial_{\mu} \psi + \mathbf{v} \partial_{\mu} \chi$$
  
 $\partial^{\mu} j_{\mu}(\mathbf{x}) = 0$ 

It is the minimum energy configuration which is not invariant.

 Goldstone Theorem: Spontaneous breaking of a continuous symmetry ⇒ A massless particle (Needs Lorentz invariance and positivity)

Consider the gauge theory extension of the previous model:

$$\mathcal{L}_1 = -rac{1}{4} \mathcal{F}_{\mu
u}^2 + |(\partial_\mu + \textit{ieA}_\mu)\phi|^2 - M^2 \phi \phi^* - \lambda (\phi \phi^*)^2$$

 $\mathcal{L}_1$  is invariant under the gauge transformation:

$$\phi(x) \rightarrow e^{i\theta(x)}\phi(x)$$
 ;  $A_{\mu} \rightarrow A_{\mu} - \frac{1}{e}\partial_{\mu}\theta(x)$ 

・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・
 ・

Consider the gauge theory extension of the previous model:

 *L*<sub>1</sub> = -<sup>1</sup>/<sub>4</sub>*F*<sup>2</sup><sub>µν</sub> + |(∂<sub>µ</sub> + ieA<sub>µ</sub>)φ|<sup>2</sup> - M<sup>2</sup>φφ<sup>\*</sup> - λ(φφ<sup>\*</sup>)<sup>2</sup>

 *L*<sub>1</sub> is invariant under the gauge transformation:
 φ(x) → e<sup>iθ(x)</sup>φ(x) ; A<sub>µ</sub> → A<sub>µ</sub> - <sup>1</sup>/<sub>e</sub>∂<sub>µ</sub>θ(x)

 Same analysis for λ > 0 and M<sup>2</sup> < 0 yields:
 </li>

$$\mathcal{L}_{1} \rightarrow \mathcal{L}_{2} = -\frac{1}{4}F_{\mu\nu}^{2} + \frac{e^{2}v^{2}}{2}A_{\mu}^{2} + evA_{\mu}\partial^{\mu}\chi \\ + \frac{1}{2}(\partial_{\mu}\psi)^{2} + \frac{1}{2}(\partial_{\mu}\chi)^{2} - \frac{1}{2}(2\lambda v^{2})\psi^{2} + \dots$$

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ の へ ()

•  $\mathcal{L}_2$  is invariant under the gauge transformation:

$$\begin{split} \psi(x) &\to \cos\theta(x)[\psi(x) + v] - \sin\theta(x)\chi(x) - v \\ \chi(x) &\to \cos\theta(x)\chi(x) + \sin\theta(x)[\psi(x) + v] \\ A_{\mu} &\to A_{\mu} - \frac{1}{e}\partial_{\mu}\theta(x) \end{split}$$

▲□▶ ▲圖▶ ▲ 臣▶ ▲ 臣▶ ― 臣 … のへぐ

•  $\mathcal{L}_2$  is invariant under the gauge transformation:

$$\begin{split} \psi(x) &\to \cos\theta(x)[\psi(x) + v] - \sin\theta(x)\chi(x) - v \\ \chi(x) &\to \cos\theta(x)\chi(x) + \sin\theta(x)[\psi(x) + v] \\ A_{\mu} &\to A_{\mu} - \frac{1}{e}\partial_{\mu}\theta(x) \end{split}$$

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

•  $\mathcal{L}_2$  contains a term proportional to  $A^2$ . A massive photon??

•  $\mathcal{L}_2$  is invariant under the gauge transformation:

$$\begin{split} \psi(x) &\to \cos\theta(x)[\psi(x) + v] - \sin\theta(x)\chi(x) - v \\ \chi(x) &\to \cos\theta(x)\chi(x) + \sin\theta(x)[\psi(x) + v] \\ A_{\mu} &\to A_{\mu} - \frac{1}{e}\partial_{\mu}\theta(x) \end{split}$$

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- $\mathcal{L}_2$  contains a term proportional to  $A^2$ . A massive photon??
- Degrees of freedom:

 $\mathcal{L}_1: 2+2=4$ 

 $\mathcal{L}_2$ : 2+3=5 ??

Notice the term  $evA_{\mu}\partial^{\mu}\chi$ 

SSB: Gauge Symmetries. Conclusions:

#### The Brout-Englert-Higgs Mechanism

• The vector bosons corresponding to spontaneously broken generators of a gauge group become massive.

• The corresponding Goldstone bosons decouple and disappear from the physical spectrum.

• Their degrees of freedom become the longitudinal components of the vector bosons.

• Gauge bosons corresponding to unbroken generators remain massless.

• There is always at least one physical, massive, scalar particle.



#### Robert Brout, François Englert, Peter Higgs

Brout died in 2011 and did not assist to the triumph of the theory he contributed to formulate.

(日) (同) (日) (日)

The Standard Model: The full Lagrangian

$$\mathcal{L} = -\frac{1}{4} \vec{W}_{\mu\nu} \cdot \vec{W}^{\mu\nu} - \frac{1}{4} B_{\mu\nu} B^{\mu\nu} + |D_{\mu}\Phi|^{2} - V(\Phi)$$

$$+ \sum_{i=1}^{3} \left[ \bar{\Psi}_{L}^{i} i \vec{D} \Psi_{L}^{i} + \bar{R}_{i} i \vec{D} R_{i} - G_{i} (\bar{\Psi}_{L}^{i} R_{i} \Phi + h.c.) \right]$$

$$+ \bar{Q}_{L}^{i} i \vec{D} Q_{L}^{i} + \bar{U}_{R}^{i} i \vec{D} U_{R}^{i} + \bar{D}_{R}^{i} i \vec{D} D_{R}^{i} + G_{u}^{i} (\bar{Q}_{L}^{i} U_{R}^{i} \tilde{\Phi} + h.c.) \right]$$

$$+ \sum_{i,j=1}^{3} \left[ (\bar{Q}_{L}^{i} G_{d}^{ij} D_{R}^{j} \Phi + h.c.) \right]$$

$$D_{\mu} Q_{L}^{i} = \left( \partial_{\mu} - ig \frac{\vec{\tau}}{2} \cdot \vec{W}_{\mu} - i \frac{g'}{6} B_{\mu} \right) Q_{L}^{i}$$

$$D_{\mu} U_{R}^{i} = \left( \partial_{\mu} - i \frac{2g'}{3} B_{\mu} \right) U_{R}^{i}$$

$$D_{\mu} D_{R}^{j} = \left( \partial_{\mu} + i \frac{g'}{3} B_{\mu} \right) D_{R}^{i}$$

## The Standard Model: Arbitrary parameters

- The two gauge coupling constants g and g'.
- The two parameters of the scalar potential  $\lambda$  and  $\mu^2.$
- Three Yukawa coupling constants for the three lepton families,  $G_{e,\mu, au}$ .  $(m_
  u=0).$
- Six Yukawa coupling constants for the three quark families,  $G_u^{u,c,t}$ , and  $G_d^{d,s,b}$ .
- $\bullet$  Four parameters of the KM matrix, the three angles and the phase  $\delta.$

ション ふゆ アメリア メリア しょうくしゃ

• All but two come from the scalar fields.

◆□▶ 
◆□▶ 
●□▶ 
●□▶ 
●□▶ 
●□▶ 
●□▶

• The Standard Model has 17 arbitrary parameters.

They are related to masses and coupling constants and should be determined experimentally.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

All have been measured.

• The Standard Model has 17 arbitrary parameters.

They are related to masses and coupling constants and should be determined experimentally.

ション ふゆ アメリア メリア しょうくしゃ

All have been measured.

The Model gives a large number of predictions.

► The Standard Model has 17 arbitrary parameters.

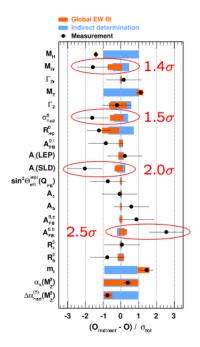
They are related to masses and coupling constants and should be determined experimentally.

All have been measured.

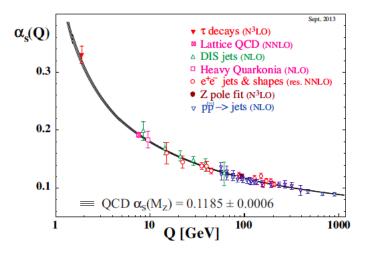
• The Model gives a large number of predictions.

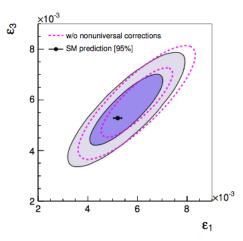
THE STANDARD MODEL HAS BEEN ENORMOUSLY SUCCESSFUL

ション ふゆ く 山 マ チャット しょうくしゃ



◆□▶ ◆圖▶ ◆臣▶ ◆臣▶ 臣 - のへで



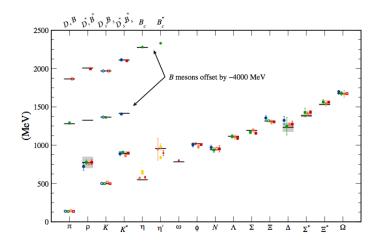


<□▶ <□▶ < □▶ < □▶ < □▶ < □ > ○ < ○

$$\epsilon_1 = \frac{3G_F m_t^2}{8\sqrt{2}\pi^2} - \frac{3G_F m_W^2}{4\sqrt{2}\pi^2} \tan^2 \theta_W \ln \frac{m_H}{m_Z} + \dots$$
(1)

$$\epsilon_3 = \frac{G_F m_W^2}{12\sqrt{2}\pi^2} \ln \frac{m_H}{m_Z} - \frac{G_F m_W^2}{6\sqrt{2}\pi^2} \ln \frac{m_t}{m_Z} + \dots$$
(2)

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ④�?



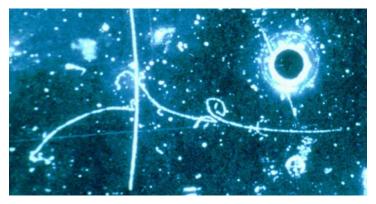
The spectrum of hadrons, computed by lattice QCD simulations and compared with the experimental results.

The precision of the measurements often led to successful predictions of new Physics.

The discovery of weak neutral currents by Gargamelle in 1972

$$u_{\mu} + e^- 
ightarrow 
u_{\mu} + e^-$$
 ;  $u_{\mu} + N 
ightarrow 
u_{\mu} + X$ 

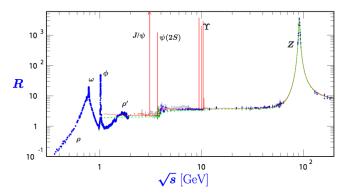
Both, their strength and their properties were predicted by the Model.



The discovery of charmed particles at SLAC in 1974

Their presence was essential to ensure the absence of strangeness changing neutral currents, ex.  $K^0\to\mu^++\mu^-$ 

Their characteristic property is to decay predominantly in strange particles.



◆□▶ ◆□▶ ◆臣▶ ◆臣▶ ─臣 ─ のへで

► A necessary condition for the consistency of the Model is that ∑<sub>i</sub> Q<sub>i</sub> = 0 inside each family.

When the  $\tau$  lepton was discovered the *b* and *t* quarks were predicted with the right electric charges.

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

A necessary condition for the consistency of the Model is that ∑<sub>i</sub> Q<sub>i</sub> = 0 inside each family.

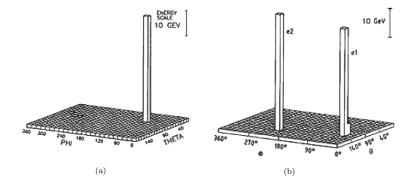
When the  $\tau$  lepton was discovered the *b* and *t* quarks were predicted with the right electric charges.

The t-quark was seen at LEP through its effects in radiative corrections before its actual discovery at Fermilab.

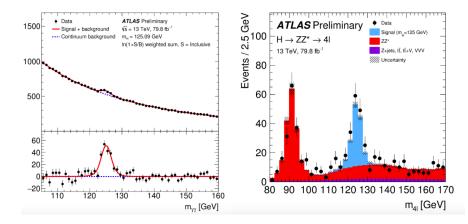
ション ふゆ アメリア メリア しょうくの

The discovery of the W and Z bosons at CERN in 1983

The characteristic relation of the Standard Model with an isodoublet BEH mechanism  $m_Z = m_W/cos\theta_W$  is checked with very high accuracy (including radiative corrections).

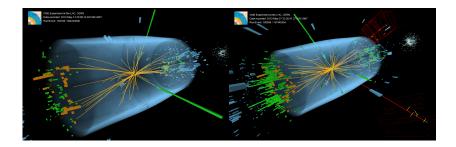


The final touch: the discovery of the BEH scalar at CERN



The discovery of the BEH scalar in the decay modes  $2\gamma$  (left) and 4/ (right). The figures include the data of  $\sqrt{s} = 13$  TeV.

#### The final touch: the discovery of the BEH scalar at CERN



Two beautiful events among those which established the discovery. The left figure shows a  $2\gamma$  decay with two photons shown as green tracks in the electromagnetic calorimeter. The right figure shows an  $e^+e^-\mu^+\mu^-$  decay with the electrons as green tracks in the e.m. calorimeter and the muons as red tracks in the muon chambers.

 Given this impressive success... What does Beyond mean?

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

- Given this impressive success... What does Beyond mean?
- Or, What is wrong with the Standard Theory??

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ ● ● ●

- Given this impressive success... What does Beyond mean?
- Or, What is wrong with the Standard Theory??

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

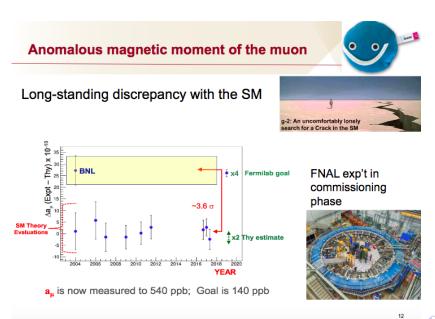
I. General questions

- Given this impressive success...
   What does Beyond mean?
- Or, What is wrong with the Standard Theory??

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

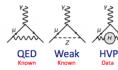
- I. General questions
- II. Specific points

High precision measurements



## High precision measurements

#### Arduous computation of ever more precise SM prediction



=

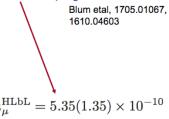
=



New lattice computation for HLBL term

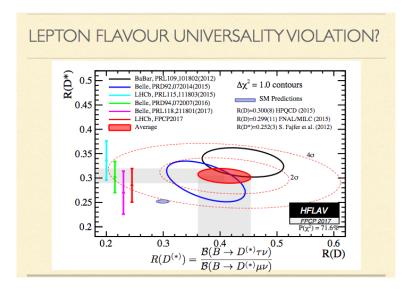
- physical pion mass and large lattice
- Statistical precision x2 improvement
- · Systematics in progress

			. \
Contribution	Value $\times 10^{10}$	Uncertainty $\times 10^{10}$	· \
QED	$11 \ 658 \ 471.895$	0.008	· \
Electroweak Corrections	15.4	0.1	
HVP (LO) [7]	692.3	4.2	
HVP (LO) [8]	694.9	4.3	
HVP (NLO)	-9.84	0.06	
HVP (NNLO)	1.24	0.01	HLbL
HLbL	10.5	2.6	$a_{\mu}^{\text{ILDL}}$
Total SM prediction [7]	11 659 181.5	4.9	. ,
Total SM prediction [8]	$11 \ 659 \ 184.1$	5.0	
BNL E821 result	11 659 209.1	6.3	•
Fermilab E989 target		$\approx 1.6$	



13

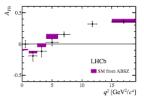
## Heavy flavour decays



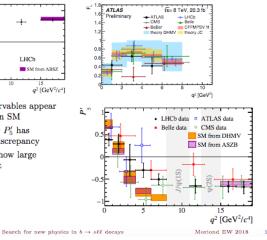
## Heavy flavour decays

Flavour changing neutral currents

 $B^0_d \to K^* \mu^+ \mu^-$  results



- Several observables appear different than SM
- In particular P'<sub>5</sub> has significant discrepancy
- Global fits show large disagreement





F. Dettori

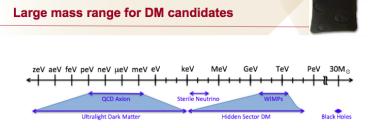
## Heavy flavour decays

Summary of B anomalies Are we there yet?

- 1. Low  $b \rightarrow s \mu \mu$  branching fractions
- 2. Discrepancies in angular observables of  $B^0_d \to K^* \mu^+ \mu^-$
- 3. Signs of lepton non-universality in:  $B^+ \to K^+ \mu^+ \mu^-$  and  $B^0_d \to K^* \mu^+ \mu^-$
- All seems to be related to a change in the C<sub>9</sub> coefficient (or maybe C<sub>9</sub> and C<sub>10</sub>, but V-A)
- Global fits start to exhibit several standard deviations of discrepancy
- $c\bar{c}$  interference explanation seems not justified
- Additional discrepancies in tree-level  $B \to D^{(*)} \ell \nu$  decays
- Many NP explanations: Z', leptoquarks, low mass resonances etc

LIVERPOOL

#### Dark matter



- bosonic DM produced during inflation or high temp phase transition
- DM acts as oscillating classical field
- WIMPs: act through SM forces
- Hidden Sector: act through new force, very weakly coupled to SM
- · Thermal contact in early universe

#### Beyond WIMPS: novel, low-cost, search techniques

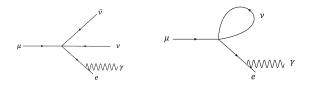
US Cosmic Visions Report, 1707.04591 23

・ロト ・ 御 ト ・ ヨ ト ・ ヨ ト ・ ヨ ・

#### For years we thought that:

• There are three distinct neutrinos.

(i) The absence of μ → e + γ shows that e.m. interactions conserve lepton numbers.
(ii) Schwinger postulated ν<sub>μ</sub> ≠ ν<sub>e</sub>.
(iii) Feinberg made it more precise:



• LEP confirmed that there are three distinct light neutrinos.

#### We also thought that all neutrinos were massless

• A totally unexplained degeneracy

(i) In Nature only *CPT* related states, i.e. particle–anti-particle, are known to be degenerate. Neutrinos were the exception.
(ii) Each neutrino carried a separate lepton number, but this was only measurable dynamically.

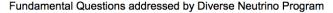
• Attempts to make at least one of the neutrinos a Goldstone fermion failed. (No Adler's decoupling.)

- The discovery that neutrinos have different masses solved this conceptual degeneracy problem.
- But created another one: Why the masses are so small?

• Neutrino physics as a portal to Physics Beyond the Standard Model.

**Neutrino Physics** 





- What is the origin of neutrino mass?
- How are the neutrino masses ordered?
  - · Oscillation experiments
- What is the absolute neutrino mass scale?
  - Beta-decay spectrum
  - Cosmic surveys
- Do neutrinos and anti-neutrinos oscillate differently?
  - · Oscillation experiments
- Are there additional neutrino types and interactions?
  - Oscillation experiments
  - Cosmic surveys
- Are neutrinos their own anti-particles?
  - Neutrinoless double-beta decay



◆□▶ ◆□▶ ◆□▶ ◆□▶ ● ● ●

My conclusion :

• A data-driven subject in which theorists have not played the major role.

• Substantial improvement in precision could be expected during the coming years.

• The significance of such improvements is not easy to judge.

• So far no real illumination came from leptons to be combined with the quark sector for a more complete theory of flavour

The trouble is that I do not see how this could change!

| ◆ □ ▶ → @ ▶ → 差 ▶ → 差 → のへぐ

Why three families

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

Why so many mass scales

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$
- Why so many mass scales
- Hierarchy and fine tuning

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$
- Why so many mass scales
- Hierarchy and fine tuning

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

Unification

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$
- Why so many mass scales
- Hierarchy and fine tuning

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- Unification
- Quantum gravity

- Why three families
- Why  $U(1) \times SU(2) \times SU(3)$
- Why so many mass scales
- Hierarchy and fine tuning
- Unification
- Quantum gravity
- Many others you can add

▲ロト ▲圖ト ▲画ト ▲画ト 三国 - のへで

No coherent picture emerges

No coherent picture emerges

► We were expecting new physics to be around the corner..... But we see no corner

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

No coherent picture emerges

► We were expecting new physics to be around the corner..... But we see no corner

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

The easy answer: We need more data

- No coherent picture emerges
- ► We were expecting new physics to be around the corner..... But we see no corner
- The easy answer: We need more data
- Two problems: (i) We do not know what kind of data
   (ii) They will not come for quite a long time

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

- No coherent picture emerges
- ► We were expecting new physics to be around the corner..... But we see no corner
- The easy answer: We need more data
- Two problems: (i) We do not know what kind of data
   (ii) They will not come for quite a long time

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()

A rather frustrating problem!

- ◆ □ ▶ → 個 ▶ → 注 ▶ → 注 → のへぐ

#### The Future of Particle Physics will undoubtedly be bright, but....

◆□ > < 個 > < E > < E > E 9 < 0</p>

The Future of Particle Physics will undoubtedly be bright, but....

▲□▶ ▲圖▶ ▲臣▶ ★臣▶ ―臣 …の�?

▶ I will not learn the answer

- The Future of Particle Physics will undoubtedly be bright, but....
- I will not learn the answer
- We have a very successful Standard Theory and we will leave the problem of its completion to the younger generation.....

・ロト ・ 日 ・ ・ 日 ・ ・ 日 ・ ・ つ へ ()